

# DS 9: The Density of States

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## 1 Density of states in $d$ –dimensions

- (a) Consider a gas of fermions (say, electrons) in a  $d$ –dimensional box, and compute the momentum-space density of states,  $g(k)dk$ .
- (b) First consider the case of non-relativistic matter. Using the dispersion relation

$$\epsilon = \frac{p^2}{2m} = \frac{\hbar^2 k^2}{2m}, \quad (1)$$

compute the density of states in energy  $g(\epsilon)$ .

*Hint:* You will need to use the fact that in  $d$ –dimensions

$$d^d k = S_d^{(1)} k^{d-1} = \frac{2\pi^{d/2}}{\Gamma(d/2)} k^{d-1}, \quad (2)$$

where  $S_d^{(1)}$  is the surface area of a unit sphere in  $d$  dimensions.

- (c) Find the Fermi energy  $\epsilon_F$  in terms of the total number of fermions  $N$ .
- (d) Use  $\epsilon_F$  to compute the internal energy  $E$  when  $T \approx 0$ .
- (e) From the variation of the internal energy with volume, compute the *pressure* of this system. This pressure, due solely to the fact that these particles are fermions, is thus known as the “degeneracy pressure”.
- (f) Now, repeat the above calculations for a gas of fermions that are *relativistic*, using the dispersion relation for relativistic particles,

$$\epsilon = pc = \hbar kc. \quad (3)$$

## 2 White Dwarfs in $d$ –dimensions

- (a) Consider a star being held at equilibrium at some radius  $R$ . The gravitational potential energy of such a system in 3-dimensions is given by

$$E_g = - \int_0^R \frac{GM(r) \times 4\pi r^2 \rho(r)}{r} dr, \quad (4)$$

where  $\rho(r)$  is the density at every point  $r$ ,  $M(r)$  is the mass contained in a radius  $r$ , and  $G$  is the gravitational constant. What would the corresponding potential energy be in  $d$  dimensions?

- (b) If we assume the density of this system to be a constant  $\rho$ , compute the integral for the gravitational energy in  $d$  dimensions.
- (c) Use this number to compute the gravitational *pressure* in  $d$  dimensions.
- (d) In a very crude sense, White Dwarfs are stable because they can find a radius at which their gravitational pressures and degeneracy pressures can equalise. Are White Dwarfs stable in all dimensions?